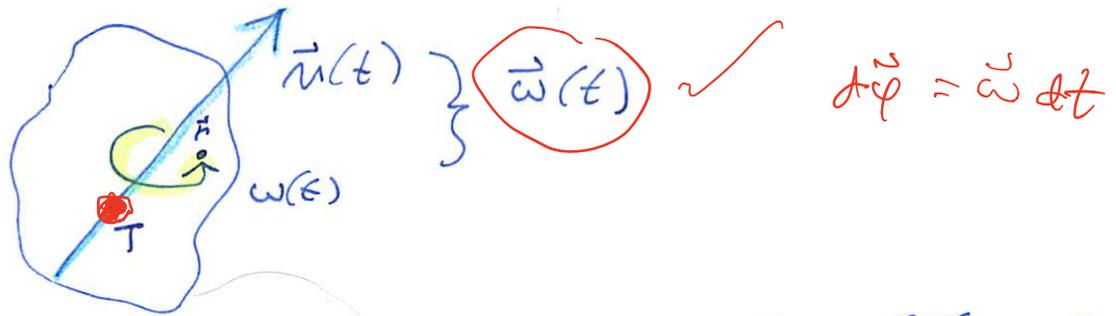


Telo z 1 nepremično (fiksnim) točko



Točko T naj bo nepremična (translacijsko samo odtransformirali). V vsakem trenutku je gibanje telesa po Eulerjevem izreku rotacija okoli neke osi (\vec{n}). Za poljubno točko \vec{r} v telesu velja

$$\frac{d\vec{r}}{dt} = \left(\frac{d\vec{r}}{dt} \right)_{\text{rel.}} + \underbrace{\vec{\omega} \times \vec{r}} dt \quad \vec{\omega} dt = d\vec{\varphi}(t)$$

hitrost, ker je v neinercialnem sistemu; $\vec{a}_{\text{rel.}} = 0$, ker je telo togo in miruje glede na sistem, ki je prifet na telo (trivialno).

keraj

$$\vec{v}(t) = \vec{\omega}(t) \times \vec{r}(t) \quad ; \quad \text{glede na } \text{inerc. sistem (holonotajzmen)}$$

($\vec{R} = \text{konst.}$)

Utilna količina telesa (pozornost)

$$\vec{L} = \sum_{i=1}^N \vec{l}_i = \sum_i m_i (\vec{r}_i \times \vec{v}_i) = \sum_i m_i (\vec{r}_i \times (\vec{\omega} \times \vec{r}_i))$$

($\vec{\omega}$ v nehem trenutku)

$$\vec{L} = \sum_i m_i [(\vec{r}_i \cdot \vec{r}_i) \vec{\omega} - (\vec{r}_i \cdot \vec{\omega}) \vec{r}_i] = \underline{\underline{J}} \vec{\omega}$$

rotacijski moment

Pogosto je ugodno uvesti gostotes masi,

$$\rho(\vec{r}) = \lim_{\Delta V \rightarrow 0} \frac{\Delta m}{\Delta V} \sim \frac{dm}{dV}; \quad \Delta m = \sum_i m_i$$

$$M = \sum_i m_i \rightarrow \int \rho(\vec{r}) dV$$

Vrotajni moment:

$$\underline{J} = \int \rho(\vec{r}) dV \begin{pmatrix} y^2+z^2 & -xy & -xz \\ -yx & x^2+z^2 & -yz \\ -zx & -zy & x^2+y^2 \end{pmatrix}$$

realen

oz.: $\underline{J}_{\alpha\beta} = \int \rho(\vec{r}) (\delta_{\alpha\beta} x^2 - x_\alpha x_\beta) d^3r = \underline{J}_{\beta\alpha}$

simetričen tenzor

"tenzor" je matrica, ki se ob vrtenju transformira pri rotaciji sistema in je v novem sistemu znesa med lokalnimi enotami. Upr.:

$$\underline{L}' = R \underline{L} \quad \text{in} \quad \underline{\omega}' = R \underline{\omega}$$

$$\underline{L} = \underline{J} \underline{\omega} \Rightarrow \underline{L}' = \underline{J}' \underline{\omega}' \quad \underline{J}' = R \underline{J} R^{-1}$$

dolost:

$$\underline{L}' = R \underline{L} = R \underline{J} \underline{\omega} = R \underline{J} R^{-1} R \underline{\omega} = \underline{J}' \underline{\omega}'$$

J ima 3 lastne vrednosti in 3 lastne vektore;

$$\underline{J} \vec{e}_\alpha = \lambda_\alpha \vec{e}_\alpha \quad \lambda_\alpha = J_{1,2,3}; \quad J_{\alpha\alpha} \geq 0$$

če $J_\alpha \neq J_\beta \Rightarrow \vec{e}_\alpha \cdot \vec{e}_\beta = 0$, ker je J realen in simetričen.

.. Energija rotirajočega telesa, (V=0)

$$T = \frac{1}{2} \sum_i m_i v_i^2 = \frac{1}{2} \sum_i m_i (\vec{\omega} \times \vec{r}_i) \cdot \vec{v}_i =$$

$$= \frac{1}{2} \sum_i m_i \vec{\omega} \cdot (\vec{r}_i \times \vec{v}_i) = \frac{1}{2} \vec{\omega} \cdot \vec{L} =$$

$$= \frac{1}{2} \vec{\omega} \cdot \underline{J} \vec{\omega}$$

$W_r = \frac{1}{2} J \omega^2$

$\vec{e}_1 \cdot \vec{e}_3 = \delta_{13}$

V rotirajočem sistemu (rotirajoče telo, en sam "piquet" na telo) velja

$\underline{J} = \begin{pmatrix} J_1 & & \\ & J_2 & \\ & & J_3 \end{pmatrix}$ in $\vec{\omega} = \sum_{d=1}^3 \omega_d \vec{e}_d$

$\vec{L} = \underline{J} \vec{\omega} = \sum_{d=1}^3 J_{dd} \omega_d \vec{e}_d = \vec{L}$ rotirajoči sistem

$\frac{1}{2} \vec{\omega} \cdot \underline{J} \vec{\omega} = \frac{1}{2} \sum_{d=1}^3 J_{dd} \omega_d^2 \geq 0$ za $\forall \vec{\omega}$ se miči iz definicije ($x^2+z^2 \geq 0$) itd.

$\Rightarrow J_{dd} \geq 0$

rotirajoče mednosilnosti ≥ 0 .

\underline{J} je pozitivno definitna matrika.

Euročke gibanja

$\frac{d\vec{L}}{dt} = \left(\frac{d\vec{L}}{dt} \right)_{\text{rotir.}} + \vec{\omega} \times \vec{L} = \vec{M}$

glede na telo $\vec{L} \neq \text{nijsno } 0$, kot sledi.

Rotacija v rotacijskem sistemu J ,

$$\vec{L} = \underline{J} \vec{\omega} = \sum_{\alpha=1}^3 J_{\alpha\alpha} \omega_{\alpha} \vec{e}_{\alpha} = \begin{pmatrix} J_1 \omega_1 \\ J_2 \omega_2 \\ J_3 \omega_3 \end{pmatrix}; \quad J_2 \leftarrow J_{22} \text{ itd.}$$

V rotacijskem sistemu je v gornjem izrazu od časa lahko odvisen le vrhota $\vec{\omega}$; ostalo, ker $J_{\alpha\alpha}$ so itak konstante in ena je definicija \vec{e}_{α} . Torej velja

$$\rightarrow \left(\frac{d\vec{L}}{dt} \right)_{\text{inert.}} = \sum_{\alpha=1}^3 J_{\alpha\alpha} \dot{\omega}_{\alpha} \vec{e}_{\alpha}$$

? $L_{\alpha}(t)$ torej $\neq 0$ glede na lastni sistem (v ravnini)

Velja še

$$\vec{\omega} \times \vec{L} = \left((J_3 - J_2) \omega_2 \omega_3, (J_1 - J_3) \omega_3 \omega_1, (J_2 - J_1) \omega_1 \omega_2 \right)$$

$\begin{pmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{pmatrix} \times \begin{pmatrix} J_1 \omega_1 \\ J_2 \omega_2 \\ J_3 \omega_3 \end{pmatrix} = \dots$ ciklično

torej $\rightarrow 1 \rightarrow 2 \rightarrow 3$

$$\begin{cases} J_1 \dot{\omega}_1 - (J_2 - J_3) \omega_2 \omega_3 = M_1 \\ J_2 \dot{\omega}_2 - (J_3 - J_1) \omega_3 \omega_1 = M_2 \\ J_3 \dot{\omega}_3 - (J_1 - J_2) \omega_1 \omega_2 = M_3 \end{cases}$$

nelinearna
1. reda
nelinearna

To so Eulerjeve enačbe; v lastnem sistemu J .



1. red, nelinearna.

$$\vec{\omega}_T$$

Najenostavnejši primer je, ko mi zunanjih momentov, $\vec{M} = 0$, potem

$$\frac{d\vec{L}}{dt} = 0 \Rightarrow \vec{L} = \vec{L}_0 = \text{const.} \quad (\text{konstantno} \dots)$$

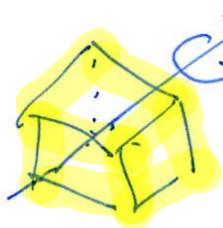
To je zapisano v inercialnem sistemu. Vzporeden mi restimo. Je pa enostavno za

$J_1 = J_2$. Če $J_1 = J_2 = J_3 = J$ je trivialno,

$\vec{L} = J \vec{\omega}$
 $J_\alpha \dot{\omega}_\alpha = J \dot{\omega}_\alpha = 0 \quad \forall \alpha \Rightarrow \vec{\omega}(t) = \vec{\omega}_0 = \text{const.}$

Pomembni: pričakujem, da las vedno telo, benda v sponnem mi. Telo simetrično inna zmeda krogla, pa tudi kocka. Če torzi v zseh vrtens kocka, ho $\vec{L} = \text{const.}$ (zmeda) in tudi $\vec{\omega} = \text{const.}$

$$\vec{\omega} \propto \vec{L}; \quad \vec{L} = (J_{ij}) \vec{\omega} = J \vec{\omega}$$



\vec{L} konst.

$$\underline{J} = J \underline{I}$$

Dolji zvrstitev je primer, ko

$$J_1 = J_2 \neq J_3$$

se vedno neje

$$J_3 \dot{\omega}_3 = 0 \Rightarrow \omega_3 = \text{const.} = \omega_0$$

(ne pozabimo: \vec{e}_3 je glede na telo)

in

$$J_1 \dot{\omega}_1 - (J_1 - J_3) \omega_2 \omega_0 = 0$$

$$J_1 \dot{\omega}_2 - (J_3 - J_1) \omega_1 \omega_0 = 0$$

/:

$$\frac{\dot{\omega}_1}{\dot{\omega}_2} = - \frac{\omega_2}{\omega_1} \Rightarrow \left. \begin{aligned} \dot{\omega}_1 &= -\Omega \omega_2 \\ \dot{\omega}_2 &= \Omega \omega_1 \end{aligned} \right\} \ddot{\omega}_1 + \Omega^2 \omega_1 = 0$$

$$\Omega = \frac{J_3 - J_1}{J_1} \omega_0$$

weiter reiten je

$$\omega_1 = A \cos(\Omega t + \delta)$$

$$\omega_2 = A \sin(\Omega t + \delta)$$

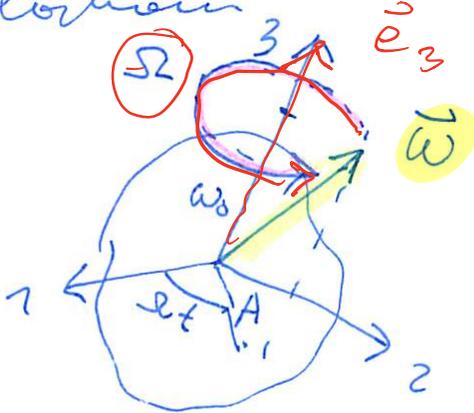
$$\omega_1^2 + \omega_2^2 = A^2$$

$$|\vec{\omega}|^2 = A^2 + \omega_0^2$$

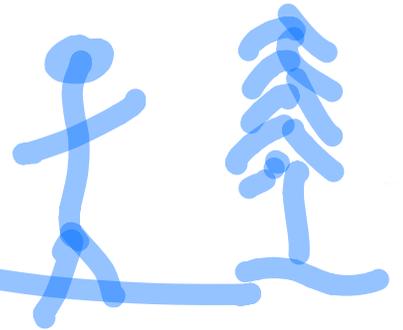
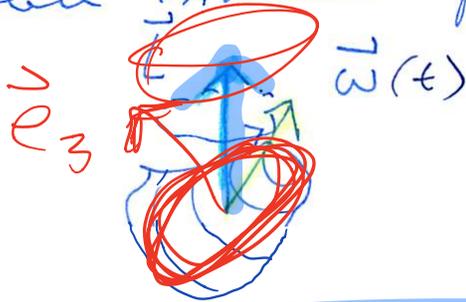
$$\vec{L} = J \vec{\omega}$$

$$\dot{\vec{\omega}} \perp \vec{\omega}$$

V lotharem System



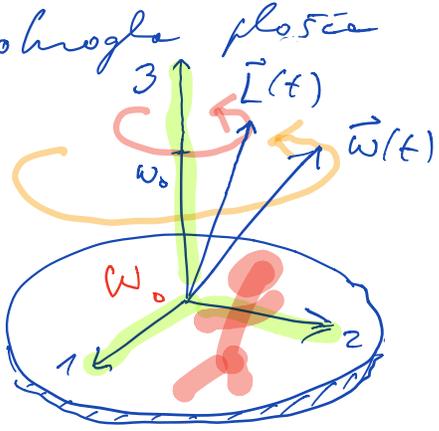
V invariante System je $\vec{L} = \text{const.}$



Primeri treserija (wobble)

Feynman
 Zgleda se solite
 Mo. F. ?

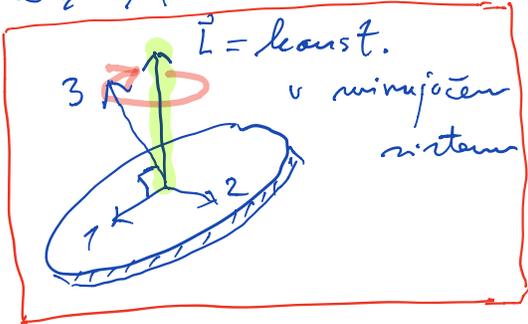
(a) Frisbee; ohropla plošča



$$\omega_3 = \omega_0$$

$$\omega_1 = A \cos \Omega t$$

$$\omega_2 = A \sin \Omega t$$



$$J_3 = \frac{1}{2} m R^2$$

$$J_1 = J_2 = \dots = \frac{1}{4} m R^2$$

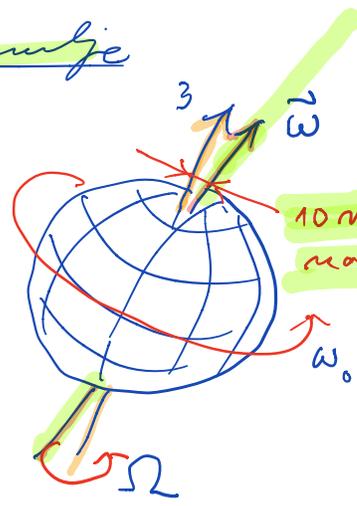
$$\Omega = \frac{J_3 - J_1}{J_1} \omega_0 = \frac{\frac{1}{2} - \frac{1}{4}}{\frac{1}{4}} \omega_0 = \omega_0 \quad (\text{v portnem sistemu})$$

$$\vec{L} = \underline{J} \vec{\omega} = \begin{pmatrix} J_1 \omega_1 \\ J_2 \omega_2 \\ J_3 \omega_3 \end{pmatrix}$$

ni || z $\vec{\omega}$ ali z \vec{e}_3

opomba: če gledamo od zornice, je $\Omega = 2\omega_0$.
 (Feynman v banu...)

(b) treserija Zemlje
 (nutacija).



10 m in se spreminja
 na polu

$$\omega_0 = 2\pi / 1 \text{ dan}$$

$$\frac{J_3 - J_1}{J_1} = 0.00327 \Rightarrow \Omega \sim 2\pi / 300 \text{ dni}$$

v resnici je 433 dni

veliko razlogov za odstopanja

- težice nutacijot
- parniki ledena na polih
- ... 27

Stabilnost vrtožbe

Naj velja redokaj

$$J_1 \neq J_2 \neq J_3 \neq J_1$$

in telo se naj vrtili okoli ene od lastnih osi. Zmagnjnih momentov ni ($\vec{M} = 0$). Eulerjeve enačbe so enačbe, v lastnem sistemu,

$$\begin{cases} J_1 \dot{\omega}_1 - (J_2 - J_3) \omega_2 \omega_3 = 0, \\ J_2 \dot{\omega}_2 - (J_3 - J_1) \omega_3 \omega_1 = 0, \\ J_3 \dot{\omega}_3 - (J_1 - J_2) \omega_1 \omega_2 = 0. \end{cases}$$

Naj se vrtili okoli osi 3,

$$\omega_1 = \omega_2 = 0 \quad \text{in} \quad \omega_3 = \Omega.$$

Kaj se zgodi, če malo spremenimo vrtenja $\vec{\omega} \rightarrow \vec{\omega} + \vec{\eta}$, $|\eta| \ll |\omega|$?

$$\omega_1 = \eta_1,$$

$$\omega_2 = \eta_2,$$

$$\omega_3 = \Omega + \eta_3.$$

Vstavimo v E. enačbe

$$J_1 \dot{\eta}_1 - (J_2 - J_3) \eta_2 (\Omega + \cancel{\eta_3}) = 0,$$

$$J_2 \dot{\eta}_2 - (J_3 - J_1) (\Omega + \cancel{\eta_3}) \eta_1 = 0,$$

$$J_3 (\dot{\Omega} + \dot{\eta}_3) - (J_1 - J_2) \cancel{\eta_1} \cancel{\eta_2} = 0.$$

Ker je $|\eta_i| \ll \Omega$ in $\dot{\Omega} = 0$, velja

$$J_1 \ddot{\eta}_1 = (J_2 - J_3) \Omega \dot{\eta}_2 + \mathcal{O}(\eta_i \eta_j)$$

$$J_2 \ddot{\eta}_2 = (J_3 - J_1) \Omega \dot{\eta}_1 + \dots \text{zanemajamo}$$

$$J_3 \ddot{\eta}_3 = 0 + \dots$$

iz tretje enačbe sledi

$$\eta_3 = \text{konst.}$$

in ko druga vstavimo v prvo, dobimo

$$J_1 \ddot{\eta}_1 = \frac{(J_2 - J_3)(J_3 - J_1)}{J_2} \Omega^2 \eta_1.$$

to je

$$c = \pm |c|$$

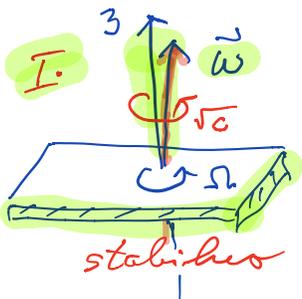
$$\ddot{\eta}_1 - c \eta_1 = 0, \quad c = \frac{(J_2 - J_3)(J_3 - J_1)}{J_1 J_2} \Omega^2.$$

Rešitve:

(a) $c < 0$: $\eta_1 = A \cos(\sqrt{|c|}t + \delta)$ ✓
 $\eta_2 = \frac{J_1}{\Omega(J_2 - J_3)} \dot{\eta}_1 = B \sin(\sqrt{|c|}t + \delta)$ ✓

$$\frac{J_1}{\Omega(J_2 - J_3)} \frac{\Omega}{\sqrt{J_1 J_2}} \sqrt{|(J_2 - J_3)(J_3 - J_1)|} = \sqrt{\frac{J_1}{J_2} \left| \frac{J_3 - J_1}{J_2 - J_3} \right|} = -\frac{B}{A}$$

$c < 0$ če $J_2 > J_3$ in $J_3 < J_1$ stabilno
 ali $J_2 < J_3$ in $J_3 > J_1$ II. 



$\Rightarrow J_3 < (J_1, J_2)$ ali $J_3 > (J_1, J_2)$
 to je najmanjši ali največji in
 gibanje je stabilno in periodično

b) $c > 0$: $\eta_1 = A_{\pm} e^{\pm \sqrt{c} t}$
 $\eta_2 = B_{\pm} e^{\pm \sqrt{c} t}$

to je zbirka, ho velja

$\rightarrow (J_2 - J_3)(J_3 - J_1) > 0$, oz.

$J_2 > J_3$ in $J_3 > J_1$

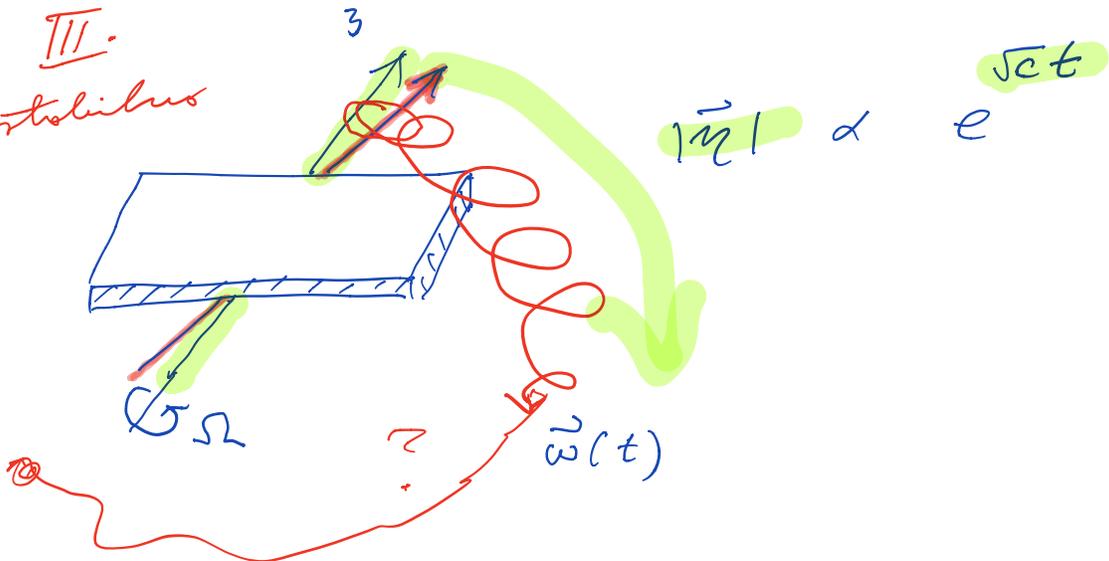
ali $J_2 < J_3$ in $J_3 < J_1$

$\Rightarrow J_1 < J_3 < J_2$ ali $J_2 < J_3 < J_1$

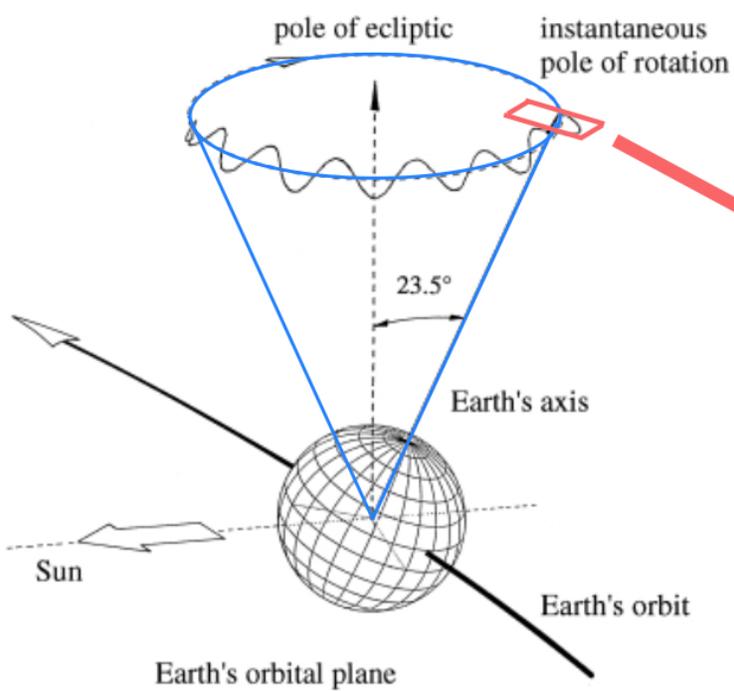
Torej je med J_1 in J_2 in
 gibanje je nestabilno

Opomba: to ne pomeni za $|\vec{\eta}| \propto e^{\pm \sqrt{c} t}$. Ko se
 amplituda poveča, η_3 d e
 linearizirane Eulerjeve enačbe
 ne veljajo več in se zapletajo

III.
 nestabilno



poizvedba celotna resitev za $\vec{\eta}(t)$ je periodična
 (npr. pojav Džamihelova).



Main cause: gravitational torque from the Moon, the Sun and the planets

